## Unit-4

## Quantum Mechanics

## Syllabus:

Black body radiation, ultraviolet catastrophe, Crompton effect, plates theory of radiation, phase and group velocity, particle in a box, uncertainty principle, well-behaved wave equation, Schrodinger equation, application to particle in a box.

## Black body:

A black body is one which absorbs all types of heat radiation incident on it when radiations are permitted to fall on black body they are neither reflected nor transmitted.
A black body is known as black body due to the fact that whatever may the colour of the incident radiation the body appears black by absorbing all kind of radiations incident on it.
A perfect black does not exists thus a body representing close proximity to perfect black body so it can be considered as a black body.
A hollow sphere is taken with fine hole and a point projection in front of the hole and is coated with lamp black on its inner surface shows the close proximity to the black body, when the radiation enter through hole, they suffer multiple reflection and are totally absorbed.

## Black Body radiation:

A body which completely absorbs radiation of all radiations of all wavelength/frequencies incident on it and emits all of them when heated at higher temperature is called black body. The radiation emitted by such a body is called black body radiation. So the radiation emitted form a black body is a continuous spectrum i.e. it contains radiation of all the frequencies.
Distributions of the radiant energy over different wavelength in the black body radiation at a given temperature are shown in the figure.
Black body radiation is a common synonym for thermal radiation.


Figure(2): Black body radiation

## Radiation:

Radiation is a process which the surface of an object radiates its thermal energy in the form of the electromagnetic waves.


Radiations are of two types

## Emissivity:

The emissivity of a material is the irradiative power of its surface to emit heat by radiation, usually it is shown by $e$ or $\varepsilon$. It is the ratio of energy radiated by a material to the energy radiated by the black body.
True black body has maximum emissivity $\varepsilon=$ (hilghly polished silver has an emissivity for about 0.02 at least.)

## Plank's Quantum Hypothesis:

Plank assumes that the atoms of the wall of blackbody behave as an oscillator and each has a characteristic frequency of oscillation. He made the following assumption-

1) An oscillator can have any arbitrary value of energy but can have only discrete energies as per the following relation

$$
E=n h v
$$

Where $n=0,1,2,3$ and $v$.and $h$ are known as frequency and Plank's constant.
2) The oscillator can absorb or emit energy only in the form of packets of energy ( $h v$ ) but not continuously.

$$
\Delta \mathrm{E}=\Delta h v
$$

## Average energy of Plank's Oscillators:

If $N$ be the total number of oscillations and $E$ as the total energy of these oscillators, then average energy will be given by the relation.

$$
\begin{equation*}
\overline{\mathrm{E}}=\frac{\mathrm{E}}{N} \tag{1}
\end{equation*}
$$

 $0, \quad h v, 2 h v \quad .$. respectikely. Then by the Maxwell's distribution formula

$$
\begin{align*}
& N=N_{0}\left(1+-\frac{h v}{k \bar{E}}+-\frac{2 h v}{e^{k T}}+\cdots \quad \ldots \quad . . \quad .\right) \\
& N=\frac{N_{0}}{\left(1--\frac{h v}{k \bar{E}}\right)} \tag{2}
\end{align*}
$$

And the total energy

$$
\begin{align*}
& E=\left(N_{0} \times\right) 0+\left(N_{1} \times h\right) v+\left(N_{2} \times 2 h\right) v+\cdots \quad \ldots . \\
& \left.E=\left(N_{0} \times\right) 0+\left(\delta e^{-\frac{h v}{k T}} \times h \nu\right){ }_{0} e^{-\frac{2 h v}{f T N}} \times 2 h v\right)+ \\
& E=N_{0} e^{-\frac{h v}{k T}} \times\left[h v+-\frac{h v}{2 \bar{E}}+3^{-\frac{2 h v}{k T}}+\cdots\right] . \\
& E=N_{0} e^{-\frac{h v}{k T}} \frac{h v}{\left(1--\frac{h v}{k \mathscr{E}}\right)^{2}} \tag{3}
\end{align*}
$$

Putting the value of $N$ and $E$ from above equations in equation (1) we get-

$$
\begin{align*}
\bar{E} & =\frac{E}{N} \\
\bar{E} & \left.=\frac{h v e^{-\frac{h v}{k T}}}{\left(1-\frac{h v}{k \bar{B}}\right.}\right) \\
\bar{E} & =\frac{h v}{\left(e^{\frac{h v}{k T}}-1\right)} \tag{4}
\end{align*}
$$

This is the expression for the average energy in Plank's oscillators.

## Plank's radiation formula:

The average density of radiation $\left(u_{v}\right)$ in the frequency range $v$ and $v+$ dapending upon the average of an oscillator is given by-

$$
\begin{align*}
& u_{v} d v=\frac{8 \pi v^{2}}{c^{3}} d v \times \overline{\times} E  \tag{5}\\
& u_{v} d v=\frac{8 \pi v^{2}}{c^{3}} \frac{h v}{\left(e^{\frac{h v}{k T}}-1\right)^{2}} d v \\
& u_{v} d v=\frac{8 \pi h}{c^{3}} \frac{v^{3}}{\left(e^{\frac{h v}{k T}}-1\right)^{2}} d v \tag{6}
\end{align*}
$$

The above relation is known as the Plank's radiation formula in terms of the frequency. This law can also be expressed in terms of wavelength $\lambda$ of the radiation. Since $v \frac{\overline{\bar{\lambda}}}{\bar{\lambda}}$ for electromagnetic radiation, $d v=$ $-\frac{c}{\lambda^{2}} d \lambda$. Further we know that the frequency is reciprocal of wavelength or in other words an increase in frequency corresponds to a decrease in wavelength. therefore

$$
\begin{align*}
& u_{\lambda} d \lambda=-u_{\nu} d \nu \\
& u_{\lambda} d \lambda=-\frac{8 \pi h}{c^{3}} \frac{\left(\frac{c}{\lambda}\right)^{3}\left(-\frac{c}{\lambda^{2}} d \lambda\right)}{\left(e^{\frac{h c}{\lambda k T}}-1\right)} \\
& u_{\lambda} d \lambda=\frac{8 \pi h c}{\lambda^{5}} \frac{1}{\left(e^{\frac{h c}{\lambda k T}}-1\right)} d \lambda \tag{7}
\end{align*}
$$

The above relation is known as the Plank's law in terms of wavelength $(\lambda)$

## Wien's law and Rayleigh-Jeans law:

With the help of Plank's radiation Wien's law and Rayleigh-Jens law can be derive. When the wavelength ( $\lambda$ ) and temperature $(T)$ are very small, then $e^{\frac{h c}{\lambda k T}} \gg$.1Therefore, 1 can be neglected in the denominator of equation (7).

$$
u_{\lambda} d \lambda=\frac{8 \pi h c}{\lambda^{5}} e^{-\frac{h c}{\lambda k T}} d \lambda
$$

By substituting $8 \pi h c=$ andd $\frac{h c}{k}=B$ we get

$$
\begin{equation*}
u_{\lambda} d \lambda=\frac{A}{\lambda^{5}} e^{-\frac{B}{k T}} d \lambda \tag{8}
\end{equation*}
$$

This is known as Wien's law, which is valid at low temperature $T$ and small wavelength $\lambda$.
For higher temperature $T$ and large wavelength $\lambda, e^{\frac{h c}{\lambda k T}}$ can be approximated to $1 \quad \frac{h c}{\lambda k T}$. Then we have from equation (7)

$$
\begin{align*}
& u_{\lambda} d \lambda=\frac{8 \pi h c}{\lambda^{5}\left(1 \frac{h c}{\lambda k T}-1\right)} d \lambda \\
& u_{\lambda} d \lambda=\frac{8 \pi k T}{\lambda^{4}} d \lambda \tag{9}
\end{align*}
$$

This is known as Rayleigh-Jeans law.

## Ultraviolet Catastrophe:

One of the nagging questions at the time concerned the spectrum of radiation emitted by a so-called black body. A perfect black body is an object that absorbs all radiation that is incident on it. Perfect absorbers are also perfect emitters of radiation, in the sense that heating the black body to a particular temperature causes the black body to emit radiation with a spectrum that is characteristic of that temperature. Examples of black bodies include the Sun and other stars, light bulb filaments, and the element in a toaster. The colours of these objects correspond to the temperature of the object. Examples of the spectra emitted by objects at narticular temneratures are shown in Figure 3


Figure 3: The spectra of electromagnetic radiation emitted by hot objects. Each spectrum corresponds to a particular temperature. The black curve(dotted line) represents the predicted spectrum of a 5000 K black body, according to the classical theory of black bodies

At the end of the 19th century, the puzzle regarding blackbody radiation was that the theory regarding how hot objects radiate energy predicted that an infinite amount of energy is emitted at small wavelengths, which clearly makes no sense from the perspective of energy conservation. Because small wavelengths correspond to the ultraviolet end of the spectrum, this puzzle was known as the ultraviolet catastrophe. Figure 27.1 shows the issue, comparing the theoretical predictions to the actual spectrum for an object at a temperature of 5000 K . There is clearly a substantial disagreement between the curves

## Matter wave:

According to Louis de-Broglie every moving matter particle is surrounded by a wave whose wavelength depends up on the mass of the particle and its velocity. These waves are known as matter wave or deBroglie waves.

## Wavelength of the de-Broglie wave:

Consider a photon whose energy is given by $E=h \gamma^{h c}=\quad[\because \quad c=\ldots . . \vartheta \cdot \lambda$
Where $h$ is Plank's constan $623 \times$ - 解 $J \sec , \vartheta$ is the frequency and $\lambda$ is the wavelength of photon. Now by Einstein's mass energy relation

$$
\begin{equation*}
E=m c^{2} \tag{2}
\end{equation*}
$$

By equation (1) and (2)

$$
\begin{aligned}
m c^{2} & =\frac{h c}{\lambda} \\
\lambda & =\frac{h}{m c}
\end{aligned}
$$

$$
\lambda=\frac{h}{p} \quad \text { Where } p=r
$$

In place of the photon a material particle of mass $m$ is moving with velocity $v$ then

$$
\begin{equation*}
\lambda=\frac{h}{m v} \tag{3}
\end{equation*}
$$

(i)

Now we know that the kinetic energy of the material particle of mass $m$ moving with velocity $\vartheta$ is given by-

$$
\begin{array}{rlr}
E & =\frac{1}{2} m v^{2} \\
E & =\frac{m^{2} v^{2}}{2 m} & \\
E & =\frac{p^{2}}{2 m} & {[\because \quad p=n} \\
p & =\sqrt{2 m E} &
\end{array}
$$

So by equation (3)

$$
\lambda=\frac{h}{\sqrt{2 m E}}
$$

(ii)

According to kinetic theory of gasses the average kinetic energy of the material particle is given by $E=$ $\frac{3}{2} K T$ where $K=1.38-\chi^{3} J \not \subset Q$ i.e. Boltzmann constant

$$
\begin{aligned}
\frac{1}{2} m v^{2} & =\frac{3}{2} K T \\
m^{2} v^{2} & =3 m K T \\
P^{2} & =3 m K T \\
p & =\sqrt{3 m K T}
\end{aligned}
$$

$$
\because \quad E \frac{1}{2} m c^{2}
$$

So by equation (3)

$$
\begin{equation*}
\lambda=\frac{h}{\sqrt{3 m K T}} \tag{4}
\end{equation*}
$$

## Group or Envelope of the wave:

When a mass particle moves with some velocity than it emits the matter waves, those waves interacts each other and where there they interfere constructively they form an envelope around the particle which is known as wave group or simply envelope.


Figure(2): Formation of the wave packet

## Group velocity:

Group velocity of a wave is the velocity with which the overall shape of the wave's amplitudes (modulation or envelope) of the wave propagates through space. It is denoted by $v_{g}$.

## Phase velocity:

The phase velocity of a wave is the rate at which the phase or the wave propagates in the space. It is denoted by $v_{p}$.

## Expression for Group velocity and phase velocity:

Let us suppose that the wave group arises from the combination of two waves that have some amplitude $A$ but differ by an amount $\Delta \omega$ in angular frequency and an mount $\Delta k$ in wave number.

$$
\begin{align*}
& \left.y_{1}=A \cos \Delta t-\right) k x  \tag{1}\\
& y_{2}=A \operatorname{cd}(\omega t+) \nmid t \omega(k+) \Delta x k \tag{2}
\end{align*}
$$

By the principle of superposition

$$
\begin{align*}
& y=y_{1}+2 y  \tag{3}\\
& y=A[\cos (\omega t-) k t c \cos (\omega t+) t \omega+k+) \Delta k]
\end{align*}
$$

Using the identity

$$
\left.\cos A+\cos B \frac{A}{2} \frac{+}{2}\right) \stackrel{B}{\operatorname{cossos}\left(\frac{-}{2}\right)^{B}}
$$

And $\cos (-\theta)=\cos \quad \theta$

$$
\begin{aligned}
& A+B \stackrel{2 \omega t}{=}-2 k x+\Delta \omega t-\Delta k x \quad A-B \underline{\omega t-k x-\omega t-\Delta \omega t}+\quad+\quad k \\
& A+B \frac{2 \omega t+\Delta \omega t-2 k x-\Delta k x}{=} \quad A-B \frac{-\Delta \omega t+\Delta k x}{2} \\
& A+B \frac{(2 \omega+\Delta \omega) t-(2 k+\Delta k) x}{2} \quad A-B=\frac{\Delta \omega t-\Delta k x}{2}
\end{aligned}
$$

$$
\left.y=A\left[\left\{\frac{(2 \omega+) \Delta \omega+2 k+) \Delta x}{2}\right\} \cdot \cos \frac{\Delta \omega t-4}{2}\right\}\right]^{\cos x}
$$

Let $2 \omega+\Delta \omega$ and $2 d k+\Delta k=2 k$
So we have

$$
\begin{aligned}
& y=2 A\left[\operatorname { c o s } \left(\frac{2\left(\omega t-8 k x \operatorname{ses}-\left(\frac{\Delta \omega t-2}{2}\right) \cdot k x\right.}{}\right.\right. \\
& \Rightarrow \quad y=2 A\left[\begin{array}{ll}
\cos (\omega t & \left.-) k x \cos \frac{\Delta \omega}{2}\left(t \quad \frac{\Delta k}{2} x\right)\right]
\end{array}\right.
\end{aligned}
$$

This is the resultant wave equation of superposition of two waves having the amplitude
$2 A \quad \cos \frac{\Delta \omega}{2}\left(t \quad \frac{\Delta k}{2} x\right)$ and phase $\cos (\omega t \quad-)$ kxchere $\omega$ and $k$ are mean values of angular frequency and prapogation constant of the wave.

## Phase velocity:

Since phase

$$
\omega t-1=\text { constant }
$$

Differentiating with respect to $t$ we get

$$
\begin{aligned}
\omega d t-k_{1} & =0 \\
\frac{d x}{d t} & =\frac{\omega}{k}
\end{aligned}
$$

But $v_{p}=\frac{d x}{d t}$

$$
\begin{equation*}
\Rightarrow \quad v_{p}=\frac{d x}{d t}=\frac{\omega}{k} \tag{5}
\end{equation*}
$$

## Group Velocity:

$$
\left.\begin{array}{rlrl} 
& \Rightarrow & \frac{\Delta \omega}{2} d t & \frac{\Delta k}{2} d x
\end{array}\right)=0
$$

So the group velocity

$$
\begin{equation*}
v_{g}=\frac{d x}{d t}=\frac{d \omega}{d k} \tag{6}
\end{equation*}
$$

## Relation between Group velocity and phase velocity

## 1. For dispersive and non-dispersive medium:

But by equation (5) i.e. $v_{p}=\frac{\omega}{k} \Rightarrow \omega={ }_{p} k v$
Putting into equation (6) we get

$$
\begin{aligned}
& \Rightarrow \quad v_{g}=\frac{d\left(k v_{p}\right)}{d k} \\
& \Rightarrow \quad v_{g}=v_{p} .1 \frac{d v_{p}}{d k} \\
& \Rightarrow \quad v_{g}=v_{p}+\frac{2 \pi x}{\lambda}\left(\frac{d v_{p}}{\left.d \frac{2 \pi}{\lambda}\right)}\right. \\
& \left.\Rightarrow \quad v_{g}=v_{p}+\frac{2 \pi x}{\lambda}\right) \frac{d v_{p}}{2 \pi d\left(\lambda^{-1}\right)} \\
& \left.\Rightarrow \quad v_{g}=v_{p}+\frac{1}{\lambda}\right)\left(\frac{d v_{p}}{\left(-\lambda^{-2}\right) .} d \lambda\right. \\
& \Rightarrow \quad v_{g}=v_{p}-\frac{d v_{p}}{\lambda}
\end{aligned}
$$

Different cases:

1) If $\frac{d v_{p}}{d \lambda}=0$ e. if the phase velocity does not depends on the wavelength then $v_{g}=p$ such a medium is called the non dispersive medium.
2) If $\frac{d v_{p}}{d \lambda} \neq 0$ e. if it has positive values then $v_{g}<p_{3}$ then such a medium is called the dispersive medium.

## 2. Relativistic particle:

Let us consider a de-Broglie wave associated with a moving particle of rest mass $m_{0}$ and velocity $v$, then the $\omega$ and $k$ will be given by

$$
\begin{align*}
& \Rightarrow \quad \omega=2 \pi \vartheta \\
& \Rightarrow \quad \omega=\frac{2 \pi m c^{2}}{h} \\
& \because \quad \vartheta \frac{m c^{2}}{\overline{\bar{h}}} \\
& \Rightarrow \quad \omega=\frac{2 \pi c^{2}}{h} \cdot \frac{m_{0}}{\sqrt{1 \frac{v^{2}}{c^{2}}}} \tag{8}
\end{align*}
$$

And

$$
\begin{array}{ll}
\Rightarrow & k=\frac{2 \pi}{\lambda} \\
\Rightarrow & k=\frac{2 \pi}{\left(\frac{h}{m v}\right)}
\end{array}
$$

$$
\Longrightarrow \quad k=\frac{2 \pi m_{0} v}{h \sqrt{1-\frac{v^{2}}{c^{2}}}}
$$

$\because \quad m \frac{\underline{\underline{m}}_{0}}{\sqrt{1-\frac{v^{2}}{c^{2}}}}$

Now phase velocity $v_{p}=\frac{\omega}{k}$
So

$$
\begin{align*}
& v_{p}=\frac{\left(\frac{2 \pi c^{2}}{h} \cdot \frac{m_{0}}{\sqrt{1-\frac{v^{2}}{c^{2}}}}\right)}{\left(\frac{2 \pi m_{0} v}{\sqrt[h]{1-\frac{v^{2}}{c^{2}}}}\right)} \\
& \left.v_{p}=\left(\frac{2 \pi c^{2}}{h} \cdot \frac{m_{0}}{\sqrt{1-\frac{v^{2}}{c^{2}}}}\right) \times \frac{h \sqrt{\left(1-\frac{v^{2}}{c^{2}}\right.}}{2 \pi\left(m_{0} v\right.}\right)
\end{align*}
$$

Now group velocity $v_{g}=\frac{d \omega}{d k}$
The expression can be written as

$$
\begin{equation*}
v_{g}=\frac{\left(\frac{d \omega}{d v}\right)}{\left(\frac{d k}{d v}\right)} \tag{11}
\end{equation*}
$$

In order to find the value of $v_{g}$ we have to solve the following terms-

$$
\begin{array}{ll}
\Rightarrow & \left.\begin{array}{rl}
\frac{d \omega}{d v} & =\frac{d}{d v}\left[\frac{2 \pi c^{2}}{h} \cdot \frac{m_{0}}{\sqrt{1} \frac{v^{2}}{c^{2}}}\right.
\end{array}\right] \\
\Rightarrow \quad \frac{d \omega}{d v} & =\frac{2 \pi c^{2} m_{0}}{h} \frac{d}{d v}\left(\sqrt{1 \frac{v^{2}}{c^{2}}}\right)^{-\frac{1}{2}} \\
\Rightarrow \quad \frac{d \omega}{d v} & =\frac{2 \pi c^{2} m_{0}}{h}\left(-\frac{1}{2}\right)\left(1 \frac{v^{2}}{c^{2}}\right)^{-\frac{3}{2}} \cdot\left(\frac{2 v}{c^{2}}\right) \\
\Rightarrow \quad \frac{d \omega}{d v} & =\frac{2 \pi c^{2} m_{0}}{h}\left(\frac{v}{c^{2}}\right)\left(1 \frac{v^{2}}{c^{2}}\right)^{-\frac{3}{2}}
\end{array}
$$

Again

$$
\begin{aligned}
& \Rightarrow \quad \frac{d k}{d v}=\frac{d}{d v}\left[\frac{2 \pi m_{0} v}{h \sqrt{1} \frac{v^{2}}{c^{2}}}\right] \\
& \Rightarrow \quad \frac{d k}{d v}=\frac{2 \pi m_{0}}{h}\left[\frac{d}{d v} \frac{v}{\sqrt{1 \frac{v^{2}}{c^{2}}}}\right]
\end{aligned}
$$

$$
\left.\begin{array}{l}
\Rightarrow \quad \frac{d k}{d v}=\frac{2 \pi m_{0}}{h} \cdot\left[\begin{array}{lll}
\frac{\sqrt{1} \frac{v^{2}}{c^{2}}}{l} & 1 & \frac{1}{2}\left(\text { u. } \frac{v^{2}}{c^{2}}\right)^{-\frac{1}{2}} \cdot\left(\frac{2 v}{c^{2}}\right) \\
\left(1 \frac{v^{2}}{c^{2}}\right)
\end{array}\right] \\
\Rightarrow \quad \frac{d k}{d v}
\end{array}\right] \frac{2 \pi m_{0}}{h} \cdot\left[\frac{\sqrt{1} \frac{\frac{v^{2}}{c^{2}}}{} \frac{v^{2}}{c^{2}} \cdot \frac{1}{\sqrt{1 \frac{v^{2}}{c^{2}}}}}{\left(\begin{array}{ll}
1 & \frac{v^{2}}{c^{2}}
\end{array}\right)}\right]
$$

Putting the value from (13) and (14) into (11) we get

$$
\begin{align*}
& \Rightarrow v_{g}=\left[\frac{\left\{\frac{2 \pi c^{2} m_{0}}{h}\left(\frac{v}{c^{2}}\right)\left(1 \frac{v^{2}}{c^{2}}\right)^{-\frac{3}{2}}\right\}}{\left\{\frac{2 \pi m_{0}}{h}\left(1 \frac{v^{2}}{c^{2}}\right)^{-\frac{3}{2}}\right\}}\right] \\
& \text { Group velocity } \quad v_{g}=v \tag{14}
\end{align*}
$$

By equation (10) we have

$$
v_{p .} \quad \imath=c^{2}
$$

By equation (14) i.e. $v_{g}=v$
So

$$
v_{p} \cdot g^{v}=c^{2}
$$

## 3. Non-Relativistic free Particle:

According to the de-Broglie hypothesis

$$
\lambda=\frac{h}{m v_{g}}
$$

Now the total energy

$$
\begin{align*}
E & =\frac{1}{2} m v_{g}^{2}  \tag{1}\\
E & =h \vartheta
\end{align*}
$$

$\qquad$

By equation (1) and (2)

$$
h \vartheta=\frac{1}{2} m v_{g}^{2}
$$

$$
h \vartheta=\frac{m v_{g}^{2}}{2 h}
$$

And phase velocity $v_{p}=\vartheta \lambda$
So we have

$$
\begin{aligned}
& v_{p}=\frac{m v_{g}^{2}}{2 h} \times \frac{h}{m v_{g}} \\
& v_{p}=\frac{v_{g}}{2}
\end{aligned}
$$

## Uncertainty Principle:

It is impossible to determine the exact position and momentum of a particle simultaneously. Let us consider a particle surrounded by a wave group of de-Broglie wave as shown in the figure


Figure(3): particle surrounded by a wave packet
Let us consider two such waves of angular frequency $\omega_{1}$ and $\omega_{2}$ and prapogation constant $k_{1}$ and $k_{2}$ traviling along the same direction are-

$$
\begin{align*}
& \psi_{1}=A \quad \sin \left(4 t-{ }_{1} x\right)  \tag{1}\\
& \psi_{2}=A  \tag{2}\\
& \sin \left(6 t-{ }_{2} x\right)
\end{align*}
$$

According to the principal of superposition

$$
\begin{aligned}
& \psi=\psi_{1}+z \\
& \psi=A \sin \left(4 t-{ }_{1} x\right)+A \sin \left(\omega_{2} x\right) \\
& \psi=A\left[\sin \left(\omega_{1} t \quad-{ }_{1} x\right)+\sin \left(\omega_{2} t \quad-{ }_{2} x\right)\right] \\
& \left.\psi=A\left[2 \sin \frac{\omega_{1} t-{ }_{1} x k+{ }_{2} t \omega-{ }_{2} x k}{2}\right) \cdot \cos ^{\omega} t\left(\frac{-{ }_{1} x k-{ }_{2} t \omega+{ }_{2} x k}{2}\right)\right] \\
& \left.\because \sin c+\sin D \frac{C+}{\overline{2}}\right\}^{D} \cdot \operatorname{sinc}{ }^{C}\left(\frac{t}{2}\right)^{D}
\end{aligned}
$$

Let

So we have

$$
\psi \quad=2 A\left[\begin{array}{ll}
\sin (\omega t & \left.-) k x \cos \frac{\Delta \omega}{2}\left(t \quad \frac{\Delta k}{2} x\right)\right]
\end{array}\right.
$$

The resultant wave is plotted in the figure (4). The position of the particle cannot be given with certainty it is somewhere between the one node and the next node. So the error in the measurement of the position of the particle is therefore equal to the distance between these two nodes.

Node


Figure(4): The envelope created by the superposition of two waves.
The node is formed when

$$
\begin{array}{rlrl}
\left.\cos \frac{\Delta \omega}{2} t \frac{\Delta k}{2} x\right)= & 0 \\
\Rightarrow & & \frac{\Delta \omega}{2} t \frac{\Delta k}{2} x=(2 n+) \frac{\pi}{2} &
\end{array}
$$

Thus $x_{1}$ and $x_{2}$ represents the positions of two successive nodes, then at any instant $t$, we get-

$$
\begin{align*}
& \frac{\Delta \omega}{2} t \quad \frac{\Delta k}{2} x_{1}=(2 n \quad+) \frac{\pi}{2} 1  \tag{5}\\
& \frac{\Delta \omega}{2} t \quad \frac{\Delta k}{2} x_{2}=(2 n \quad+) \frac{\pi}{2} 3 \tag{6}
\end{align*}
$$

Now on subtracting (5) from (6) we get

$$
\begin{aligned}
\frac{\Delta \omega}{2} t \quad \frac{\Delta k}{2} x_{2}-\frac{\Delta \omega}{2} t \quad \frac{\Delta k}{2} x_{1} & =2 n \cdot \frac{\pi}{2}+\frac{3 \pi}{2}-22_{2}^{\pi} \cdot \frac{\pi}{2} \\
\frac{\Delta k}{2}\left(x_{1}-2^{2}\right) & =\pi \\
\Delta k \Delta x & =2 \pi
\end{aligned}
$$

But $k \quad \stackrel{2 \pi}{\overline{\bar{\lambda}_{m}}}$

$$
\left.\Delta \frac{2(\pi}{\lambda d_{n}}\right) \quad \Delta \lambda=2 \pi
$$

Again $\lambda_{m}=\frac{h}{m v}=\frac{h}{p}$

So

$$
\begin{aligned}
2 \pi \Delta\left\{\frac{1}{\left(\frac{h}{p}\right)}\right\} \Delta x & =2 \pi \\
\frac{1}{h} \cdot \Delta p . \quad \iota & =1 \\
\Delta p \cdot \quad \Delta x & =h \\
\Delta p \cdot \quad \Delta x & \geq \hbar
\end{aligned}
$$

Where $\hbar=$
$h / 2 \pi$

## Energy and time uncertainty principle:

Let $\Delta x$ be the width of the wave packet moving along the x -axis, let $v_{g}$ be the group velocity of the wave packet and $v_{x}$ is the particle velocity along $x$-axis. Now if the wave packet moves through $\Delta x$ in $\Delta t$ time. Since $\Delta x$ is the uncertainty in the $x$-coordinates of the particle and $\Delta t$ is the uncertainty in the time i.e. given by

$$
\begin{align*}
\Delta t & =\frac{\Delta x}{v_{g}} \\
\Delta x & =v_{g} . \Delta t \tag{1}
\end{align*}
$$

If the rest mass of the particle is $m_{0}$ then the kinetic energy is given by

$$
\begin{align*}
E & =\frac{1}{2} m_{0} v_{x}^{2} \\
E & =\frac{1}{2 m_{0}} \cdot \quad \ddot{\partial} v_{x}^{2} \\
E & =\frac{1}{2 m_{0}} \cdot\left(\not \partial v_{x}\right)^{2} \\
E & =\frac{1}{2 m_{0}} \cdot 3 p \\
E & =\frac{p_{x}^{2}}{2 m_{0}} \tag{2}
\end{align*}
$$

If $\Delta p_{x}$ and $\Delta E$ are the uncetainity in the momentum and energy respectively, then differentiating (2) we get

$$
\begin{aligned}
\Delta E & =\frac{2 p_{x} \cdot \Delta p_{x}}{2 m_{0}} \\
\Delta E & =\frac{p_{x} \cdot \Delta p_{x}}{m_{0}} \\
\Delta P_{x} & =\frac{m_{0} \Delta E}{p_{x}}
\end{aligned}
$$

But $p_{x}=\boldsymbol{z} v_{x}$
So

$$
\Delta P_{x}=\frac{m_{0}}{m_{0} v_{x}} \Delta E
$$

Now by(1) and (3)

$$
\Delta x . \quad \Delta p=v_{g} . \quad \Delta \stackrel{1}{t_{v_{x}}} \Delta E
$$

But $v_{g}={ }_{x}$ vso we have

$$
\Delta x \Delta p_{x}=\Delta t . \quad \Delta E
$$

We know that
$\Delta x \Delta p_{x} \geq \hbar$
So by (4) and (5)
$\Delta t . \quad \Delta E \geq \hbar$

## Application of uncertainty Principle:

## Determination of the position of a particle with the help of a microscope:

Let us consider the case of the measurement of the position of an electron is determined. For this the electron is illuminated with light (photon). Now the smallest distance between the two points that can be resolved by microscope is given by

$$
\begin{equation*}
\Delta x=\frac{\lambda}{2 \sin } \tag{1}
\end{equation*}
$$

From the above equation it is clear that for exactness of position determination improves with a decrease in the wavelength $(\lambda)$ of liaght. Let us imagine that we are using a $(\gamma-r$ rayn) croscope of angular aperture $2 \theta$.


Figure(5): microscope
In order to observe the electron, it is necessary that at least one photon must strike the electron and scattered inside the microscope. The scattered photon can enter in the field of view $+\theta$ to - ds shown $^{\text {in }}$ the figure.


Figure(6): Scattering of an photon by an electron


The momentum ( $p$ ) of the scattered photon is $\left(\frac{h}{\lambda}\right)$ then the momentum along the $x$-axis is $(-\quad p \operatorname{sinan} \theta)$ $(+p \sin . 巴)$ then the photon of wavelength $\lambda^{\prime}$ collied to the electron then this photon recoile the electron by giving some momentum to it.

Now the uncertainty in the momentum transfer to the electron will be $\left.\Delta p \quad \neq p \quad+_{x}\right] p-\left[\begin{array}{ll}p & -_{x}\end{array}\right] p$

$$
\begin{align*}
& \left.\Delta p=\left(\frac{h}{\lambda^{\prime}}+\frac{h}{\lambda} \sin \theta\right) \frac{h}{\lambda^{\prime}}-\frac{h}{\lambda} \sin \theta\right) \\
& \Delta p=\frac{2 h}{\lambda} \sin \theta \tag{2}
\end{align*}
$$

By equation (1) and (2)

$$
\begin{array}{ll}
\Delta p . & \Delta \lambda=\frac{\lambda}{2 \sin } \times \frac{2 h \sin \theta}{\lambda} \theta \\
\Delta p . & \Delta \lambda=h \\
\Delta p . & \Delta \lambda \geq \frac{\hbar}{2}
\end{array}
$$

Where $\hbar \frac{h}{\overline{2} \pi}$

## Diffraction of electron beam by a single slit:

Suppose a narrow beam of electron passes through a narrow single slit and produces a diffraction on the screen as shown in figure.


Figure(7): Diffraction pattern of electron beam by single slit
But the theory of Fraunhofer's diffraction at a single slit (2d $\sin \theta=$ therfif)tt minima is given by

$$
2 \Delta y \sin =\lambda
$$

In producing the diffraction pattern on the screen, all the electrons have passed through the slits but we
can-not say definitely at what place of the slit. Hence the uncertainty in determining the position of electron is equal to the width of the slit ( $\Delta y$ ) then by equation (1)

$$
\begin{equation*}
\Delta y=\frac{\lambda}{2 \sin } \theta \tag{2}
\end{equation*}
$$

Initially the electrons are moving $x$-axis and hence they have no component of momentum along $y$-axis. After diffraction on the slit, they are deviated from their initial path to form the pattern. Now they have a component $p \quad \sin A \&$.y component of momentum may be anywhere between $\left(\begin{array}{lll}p & \sin & a t h) d(-\quad p\end{array} \sin \right.$ Hance the uncertainty in $y$ component is

$$
\begin{align*}
& \Delta p_{y}=p \sin \theta-\left(\begin{array}{ll}
-p & \sin \\
\theta
\end{array}\right) \\
& \Delta p_{y}=2 p \sin \theta \\
& \left.\Delta p_{y}=2 \frac{h}{\lambda}\right) \sin \tag{3}
\end{align*}
$$

By equation (2) and (3)
$\Delta y . \quad \Delta p=\frac{\lambda}{2 \sin } \times \frac{2 h}{\lambda} \sin \theta$
$\Delta y . \quad \Delta p \geq h$
$\Delta y . \quad \Delta p \geq \frac{\hbar}{2}$

## Compton Scattering:

When a beam of monochromatic radiation of sharply define frequency incident on materials of low atomic number, the rays suffers a change in frequency on scattering. This scattered beam contains two beams one having lower frequency or greater wavelength other having the same frequency or wavelength.

The radiation of unchanged frequency in the scattered beam is known as unmodified radiation while the radiation of lower frequency or slightly higher wavelength is called a modified radiation. This phenomenon is known as Compton effect.


Figure(8): Compton Scattering
The energy and momentum

| S.N. | Quantity | Before collision | After collision |
| :---: | :---: | :---: | :---: |
| 1. | Momentum of radiation | $\frac{h \vartheta}{c}$ (Where $\vartheta$ is the frequency of radiation) | $\frac{h \vartheta^{\prime}}{c}$ |
| 2. | Energy of radiation | $\mathrm{E}=$ (MOhere $\vartheta$ is the frequency of radiation) | $\mathrm{E}={ }^{\prime} \mathrm{h} \vartheta$ |
| 3. | Momentum of electron | 0 | $m v$ |
| 4. | Energy of electron | $E={ }_{0} a^{2}$ (Where $m_{0}$ is the rest mass of the electron) | $E=3(W \text { Where } m \text { is the }$ moving mass of the electron) |

By the principle of the conservation of momentum along and perpendicular to the direction of the incidence,
we get
In x-direction

```
        momentum before collision = mometum after collision
```

$$
\begin{equation*}
h \vartheta+=\frac{h \vartheta \prime}{c} \cos \theta+m v \mathrm{c} \tag{1}
\end{equation*}
$$

In $y$-direction

$$
\begin{equation*}
0+=\frac{h \vartheta \prime}{c} \sin \theta-m v \text { si } \tag{2}
\end{equation*}
$$

By equation (1)

$$
\begin{equation*}
m v c \cos =h \vartheta-{ }^{\prime} k \theta s \quad \theta \tag{3}
\end{equation*}
$$

By equation (2)

$$
\begin{equation*}
m v c \sin =h \vartheta^{\prime} \sin \theta \tag{4}
\end{equation*}
$$

Squaring equation (3) and (4) then adding we get

Now by conservation of energy

$$
\begin{align*}
& \text { Energy before collision }
\end{aligned} \begin{aligned}
h \vartheta+{ }_{0} o^{2} & =h v^{\prime}+m^{2} c \\
h \vartheta+{ }_{0} a^{2}-h^{\prime} v & =m c^{2}  \tag{6}\\
h \vartheta-{ }^{\prime} h+9 c^{2} & =m c^{2}
\end{align*}
$$

Squaring both side

$$
\left[h \vartheta-{ }^{\prime} h+\theta \partial c^{2}\right]^{2}=m^{2} c^{4}
$$

As we know that $(a+b)^{2}+c^{2} a+{ }^{2} b+{ }^{2} c+2 a b+2 b c+2 c a$
so
$(a-b)^{2}+c^{2} a+{ }^{2} b+{ }^{2} c-2 a b-2 b c+2 c a$

$$
h^{2} \vartheta^{2}+2 \vartheta^{\prime 2}+2 c^{4}-2 \hbar \vartheta \vartheta^{\prime}-2 \hbar \vartheta^{\prime} m_{0} c^{2}+2 h \vartheta \not c^{2}=m^{2} c^{4}
$$

$$
\begin{equation*}
h^{2} \vartheta^{2}+2 \vartheta^{\prime 2}-2 \hbar \vartheta \vartheta^{\prime}+2 c^{4}+2\left(\vartheta-^{\prime}\right) \cdot \vartheta \forall c^{2}=m^{2} c^{4} \tag{7}
\end{equation*}
$$

Subtracting equation (5) from (7) we get

$$
\begin{aligned}
& \begin{array}{r}
h^{2} \vartheta^{2}+2 \vartheta^{\prime 2}-2 \hbar \vartheta \vartheta^{\prime}+2 c^{4}+2\left(\vartheta-{ }^{\prime}\right) \cdot \vartheta \forall c^{2}-\begin{array}{l}
\left\{h \vartheta^{2}+2 \vartheta^{\prime 2}-\right. \\
2 h^{2} \vartheta \vartheta^{\prime} \cos \theta^{-}
\end{array}=m^{2} c^{4}-2 v^{2} c^{2}
\end{array} \\
& h^{2} \vartheta^{2}+2 \vartheta^{\prime 2}-2^{2} \vartheta \vartheta^{\prime}+2 c^{4}+2\left(\vartheta \vartheta-{ }^{\prime}\right) . \vartheta \not x^{2}-2 \vartheta^{2}-2 \vartheta^{\prime 2} \\
& +2^{2} \vartheta \vartheta \vartheta^{\prime} \cos \theta
\end{aligned}
$$

$$
\begin{align*}
& \left.m^{2} v^{2} c^{2}\left(\sin ^{2} \phi+{ }^{2} c \phi s\right)=\left(h \vartheta-{ }^{\prime} k \theta \mathrm{~s}\right) \not\right)^{2}+\left(h \vartheta^{\prime} \sin \right) \theta^{2} \\
& m^{2} v^{2} c^{2}=h^{2} \vartheta^{2}+2 \vartheta^{\prime 2} \cos ^{2} \theta-{ }^{2} \vartheta h \theta^{\prime} \cos \theta \quad 4 \vartheta^{\prime 2} h \sin ^{2} \theta \\
& m^{2} v^{2} c^{2}=h^{2} \vartheta^{2}+2{g^{\prime 2}}^{2}\left(\cos ^{2} \theta+{ }^{2} \mathrm{~s} \theta_{1}\right)-2 \psi \vartheta \vartheta^{\prime} \cos \theta \\
& m^{2} v^{2} c^{2}=h^{2} \vartheta^{2}+2 \vartheta^{\prime 2}-2 \hbar \vartheta \vartheta^{\prime} \cos \theta \\
& h^{2} \vartheta^{2}+2 \vartheta^{\prime 2}-2 \hbar \vartheta \vartheta^{\prime} \cos \theta=m^{2} v^{2} c^{2} \tag{5}
\end{align*}
$$

But $m=\frac{m_{0}}{\sqrt{1-\frac{v^{2}}{c^{2}}}}$
So $m^{2}=\frac{m_{0}^{2}}{1-\frac{v^{2}}{c^{2}}}$

$$
m^{2}=\frac{m_{0}^{2} c^{2}}{c^{2}-v^{2}}
$$

So

$$
\begin{aligned}
& -h^{2} .2 \vartheta^{\prime} \vartheta 1-c o s+\theta 2\left(\vartheta-{ }^{\prime}\right) \not \overbrace{0} c^{2}+2 c^{4}=\frac{m_{0}^{2} c^{2}}{\left(c^{2}-v^{2}\right)} c^{2}\left(c^{2}-讠\right) \\
& -2 h^{2} \vartheta \vartheta^{\prime}(1-c 0) s+\theta 2\left(\vartheta \vartheta-{ }^{\prime}\right) \omega_{0} c^{2}+\frac{2}{2} c^{4}=m_{0}^{2} c^{4} \\
& -2 h^{2} \vartheta \vartheta^{\prime}(1-c o) s+\theta 2\left(h \vartheta-{ }^{\prime}\right) \not h_{0} c^{2}=0 \\
& 2 h\left(\vartheta-{ }^{\prime}\right) \pi r_{0} c^{2}=2 h^{2} \vartheta \vartheta^{\prime}(1-c \partial) s \theta \\
& \left(\vartheta--^{\prime}\right) \imath=\frac{2 h^{2} \vartheta \vartheta^{\prime}}{2 h m_{0} c^{2}}(1-\operatorname{co}) s \theta \\
& \frac{\left(\vartheta-\vartheta^{\prime}\right)}{\vartheta \vartheta^{\prime}}=\frac{h}{m_{0} c^{2}}(1-\cos \theta) \\
& \frac{\vartheta}{\vartheta \vartheta^{\prime}}-\frac{\vartheta^{\prime}}{\vartheta \vartheta^{\prime}}=\frac{h}{m_{0} c^{2}}(1-\cos \theta) \\
& \frac{1}{\vartheta^{\prime}}-\frac{1}{\vartheta}=\frac{h}{m_{0} c^{2}}(1-\cos
\end{aligned}
$$

Multiplying by $c$ both side

$$
\begin{aligned}
\frac{c}{\vartheta^{\prime}}-\frac{c}{\vartheta} & =\frac{h c}{m_{0} c^{2}}(1-\cos \theta) \\
\lambda^{\prime}-; & =\frac{h}{m_{0} c}(1-\cos \epsilon \ldots \\
\Delta \lambda & =\frac{h}{m_{0} c}(1-\cos \theta) \\
\Delta \lambda & =\frac{h}{m_{0} c} 2 \sin \left(\frac{\theta}{2}\right) \\
\Delta \lambda & =\frac{2 h}{m_{0} c} \sin ^{2}\left(\frac{\theta}{2}\right)
\end{aligned}
$$

Where $\Delta \lambda$ is the change in the wavelength
Equation (10) shows that

1) If $\theta=0 \Rightarrow \Delta$ e. there 0 s no scattering along the direction of incidence.
2) If $\theta \quad \stackrel{\pi}{\overline{2}} \Rightarrow \Delta \lambda \frac{h}{\underline{\underline{\underline{h}}} \mathrm{C}}=\frac{6.6 \times 10^{-34}}{9 \times 10^{-31} \times 3 \times 10^{8}}=0.2426 \AA$ this wavelength is known as Compton wavelength and it is a constant quantity.
3) If $\theta=\pi \Rightarrow \frac{2 h}{m_{0} c} \lambda=0.4852$ Åo the change in the wavelength waries in accordance to the scattering angle $\theta$ and this is shown in figure.


Figure(9): Graph between angle of incidence and wavelength
Importance of Compton effect:

1) It provides the evidence of particle nature of the electromagnetic radiation.
2) This verifies the Plank's quantum hypothesis.
3) This provides an indirect verification of the following relation $m=\frac{m_{0}}{\sqrt{1-\frac{v^{2}}{c^{2}}}}$ and $E=3 m c$

## Direction of the recoil electron:

We know that

$$
\begin{aligned}
m v c \cos & =h \vartheta-' k \theta s \theta \\
m v c \sin & =h \vartheta^{\prime} \sin \theta
\end{aligned}
$$

Dividing (4) by (3) we get

$$
\begin{aligned}
& \tan \phi=\frac{h \vartheta^{\prime} \sin \theta}{h \vartheta-{ }^{\prime} k \theta s} \theta \\
& \tan \phi=\frac{\vartheta^{\prime} \sin \theta}{\vartheta-{ }^{\prime} \cos \theta}
\end{aligned}
$$

Again by equation (8) i.e.

$$
\begin{aligned}
\frac{1}{\vartheta^{\prime}}-\frac{1}{\vartheta} & =\frac{h}{m_{0} c}(1-\cos \theta) \\
\frac{1}{\vartheta^{\prime}} & =\frac{1}{\vartheta}+\frac{h}{m_{0} c} 2 \sin ^{2}\left(\frac{\theta}{2}\right)
\end{aligned}
$$

Multiplying by $\vartheta$ we get

$$
\begin{aligned}
\frac{\vartheta}{\vartheta^{\prime}} & =1 \frac{h \vartheta}{m_{0} c} 2 \sin ^{2}\left(\frac{\theta}{2}\right) \\
\frac{1}{\vartheta^{\prime}} & =\frac{1 \frac{h \vartheta}{m_{0} c} 2 \sin ^{2}\left(\frac{\theta}{2}\right)}{\vartheta} \\
\vartheta^{\prime} & =\frac{\vartheta}{1 \frac{h \vartheta}{m_{0} c} 2 \sin ^{2}\left(\frac{\theta}{2}\right)}
\end{aligned}
$$

Then by equation (11) and (12) we have

$$
\tan \phi=\frac{\left[\frac{\vartheta}{1 \frac{h \vartheta}{\frac{h}{m_{0} c} 2 \sin ^{2}\left(\frac{\theta}{2}\right)}}\right] \cdot \sin \theta}{\vartheta\left[\frac{\vartheta}{1 \frac{h \vartheta}{m_{0} c} 2 \sin ^{2}\left(\frac{\theta}{2}\right)}\right] \cdot \cos \theta}
$$

Let $\frac{h \vartheta}{m_{0} c}=\quad \alpha$
then

$$
\begin{aligned}
& \tan \phi=\frac{\left[\begin{array}{cc}
\vartheta \sin \theta \\
1+\alpha^{2}\left(\frac{\theta}{2}+2\right)
\end{array}\right]}{\vartheta\left[1 \frac{\cos \theta}{1+\alpha^{2}\left(\frac{\sin }{2}\right)}\right]}
\end{aligned}
$$

$$
\begin{aligned}
& \tan \phi=\frac{\sin \theta}{1+\alpha^{2}\left(\frac{\theta}{2 i n}\right)-\cos \theta} \\
& \tan \phi=\frac{\sin \theta}{(1-\cos \theta)++^{2}\binom{\theta}{Z}} 2 \sin \\
& \tan \phi=\frac{\left.2 \cdot \sin \frac{\theta}{2}\right)\left(\cdot \cos _{2}^{\theta}\right)( }{2 \cdot \sin \left(\frac{\theta}{2}\right)+\alpha^{2}\left(\frac{\theta}{2}\right)}
\end{aligned}
$$

$$
\begin{aligned}
& \tan \phi=\frac{\left.2 \cdot \sin \frac{\theta}{2}\right)(\cdot \operatorname{co}}{2 \cdot \sin \left(\frac{\theta}{2}\right)(1+\alpha)}( \\
& \tan \phi=\frac{\left.\cot \frac{\theta}{2}\right)}{\left[\begin{array}{ll}
1 & +]
\end{array} \alpha\right.} \\
& \tan \phi=\frac{\cot \frac{\theta}{2}}{\left[1+\frac{h \vartheta}{m_{0} c^{2}}\right]}
\end{aligned}
$$

## Wave function and its properties:

We know that height of the water surface varies periodically in water waves, the pressure of gas varies periodically in sound waves and the electric field and magnetic field varies periodically in light waves, similarly the quantities which varies periodically in case of matter waves is called the wave function. The quantity whose variations make up the matter waves. This is represented by $\psi$. This has no direct physical significance and is not an observed quantity. However the value of wave function is related to the probability of finding the particle at a given place at a given time, wave function $\psi$ is a complex quantity i.e.

$$
\psi=A+i B
$$

Conjugate of $\psi$ is

$$
\psi^{*}=A-i B
$$

And

$$
\psi \psi^{*}=|\psi|^{2}={ }^{2} A+\quad 3
$$

$|\psi|^{2}$ at a time at a particular place is the probability of finding the particle there at that time and is known as probability density $|\psi|^{2}=\psi^{*} \psi$
Let the wave function $\psi$ is specified in $x$ direction by the wave equation $\psi=-\dot{A} \omega e^{\left(t-\frac{x}{v}\right)}$
Where $\omega=2 \bmod v=\vartheta \lambda$
So

$$
\begin{align*}
\psi & =A e^{-2 \pi \vartheta i\left(t-\frac{x}{\vartheta \lambda}\right)} \\
\psi & =A e^{-2 \pi i\left(\vartheta t-\frac{x}{\lambda}\right)} \tag{1}
\end{align*}
$$

```
As E = h\vartheta 
And \lambda 
E \stackrel{h}{\overline{2}\pi}.\quad2\pi\vartheta
\lambda }\stackrel{2\pi\cdot\frac{h}{2\pi}}{=
E=2\pi\hbar\vartheta
\lambda }\stackrel{2\pi}{=
\frac{1}{\lambda}}=\frac{p}{2\pi\hbar
```

Putting these values in (1) we get

$$
\begin{align*}
\psi & =A e^{-2 \pi i\left(\frac{E}{2 \pi \hbar} t-\frac{p}{2 \pi \hbar} x\right)} \\
\psi & =A e^{-\frac{i}{\hbar}(E t-p x)} \tag{2}
\end{align*}
$$

This is the wave equation for a free particle.

## Properties of wave function:

1) It must be finite everywhere: if $\psi$ is infinite at a particular point, then it would mean an infinitely
large probability of finding the particle at that point, which is impossible. Hence $\psi$ must have a finite or zero values at any point.
2) It must be single valued: if $\psi$ has more than one value at any point, it means that there is more than one values of probability of finding the particle at that point, which is impossible.
3) It must be continuous: For Schrodinger equation $\frac{d^{2} \psi}{d x^{2}}$ must be finite everywhere. This is possible only where $\frac{\partial \psi}{\partial x}$ has no discontinuity at any boundary where potential changes. This implies that $\psi$ too must be continuous across a boundary.
4) $\psi$ must be normalised: $\psi$ must be normalised, which means that $\psi$ must to be zero as $x \quad \rightarrow \quad$ 士' $y \rightarrow \pm \infty \rightarrow$ + + $\quad \rightarrow$ order that $\int|\psi|^{2} d v$ over all space be finite constant.
If $\int_{-\infty}^{+\infty}|\psi|^{2} d v=0$ theparticle does not exists but $|\psi|^{2}$ over all space must be finite i.e. the body exits somewhere it
$\therefore \int_{x_{1}}^{x_{2}}|\psi|^{2} d v=0, \quad \infty$, compdex are not possible.
5) Normalization: $\int_{-\infty}^{+\infty}|\psi|^{2} d v=1$

As $|\psi|^{2}=\psi^{*} \psi=\quad$ probability density $(P)$
6) Probability between the limits $x_{1}$ and $x_{2}$ : This is given by
$P_{x_{1} x_{2}}=\int_{-\infty}^{+\infty}|\psi|^{2} d x$ (one dimensional)
7) Expected values: To correlate experiment and theory we define the expectation values of any parameter

$$
\langle x\rangle=\frac{\int_{-\infty}^{+\infty} x \cdot|\psi|^{2} d x}{\int_{-\infty}^{+\infty}|\psi|^{2} d x}=\frac{\int_{-\infty}^{+\infty} \psi^{*} x \psi d x}{\int_{-\infty}^{+\infty} \psi^{*} \psi d x}
$$

If $\psi$ is a normalised wave function then

$$
\int_{-\infty}^{+\infty}|\psi|^{2} d x=1
$$

So

$$
\langle x\rangle=\int_{\infty}^{+\infty}|\psi|^{2} d x=1
$$

## Orthonormal and Orthogonal wave function:

For two wave function $\psi_{1}(x)$ and $\psi_{2}(x)$ if the condition $\int_{a}^{b} \psi_{2}^{*}(x) \psi_{1}(x) d x \quad=$ ex0sts then they are said to be orthogonal wave function. Here $\psi_{2}^{*}(x)$ is the complex conjugate of $\psi_{2}(x)$.

The normalized wave function are defined by

$$
\int_{a}^{b} \frac{*}{z}(x) \psi_{1}(x) d x=1
$$

The wave function satisfying both the conditions of normalisation and orthogonally said to be orthonormal. These two conditions simultaneously can be written as

$$
\int_{a}^{b} \text { 毅 }(x) \psi_{n}(x) d \tau=\left\{_{n}=\left\{\begin{array}{l}
=\text { flor } m=r \\
=\text { for } m \neq r
\end{array}\right.\right.
$$

## Operator:

Operator $\hat{O}$ is a mathematical rule which may applied to a function $f(x)$ which changes the function in to an other function $g(x)$.

So an operator is a rule by means of which from a given function, we can find another function for example:

$$
\frac{d}{d x} e^{a x}=a^{a} e^{x}
$$

So an operator tells us that what operation to carry out on the quantity that follows it.

## Energy Operator:

We know that the wave function is given as-

$$
\psi=A e^{-\frac{i}{\hbar}(E t-p x)}
$$

Differentiating partially with respect to $t$ we get

$$
\begin{aligned}
\frac{\partial \psi}{\partial t} & =-\frac{i E}{\hbar} A e^{-\frac{i}{\hbar}(E t-p x)} \\
\frac{\partial \psi}{\partial t} & =-\frac{i E}{\hbar} \psi \\
E \psi & =-\frac{\hbar}{i} \frac{\partial \psi}{\partial t}
\end{aligned}
$$

Hence energy operator

$$
\hat{E}=i \hbar \frac{\partial \psi}{\partial t}
$$

## Momentum Operator:

Again by wave function i.e.

$$
\psi=A e^{-\frac{i}{\hbar}(E t-p x)}
$$

Differentiating equation with respect to $x$ we get

$$
\begin{aligned}
\frac{\partial \psi}{\partial x}= & \frac{i p}{\hbar} A e^{-\frac{i}{\hbar}(E t-p x)} \\
\frac{\partial \psi}{\partial x}= & \frac{i p}{\hbar} \psi \\
\frac{\hbar}{i} \frac{\partial \psi}{\partial x}= & p \psi \\
p \psi= & -i \hbar \frac{\partial \psi}{\partial x} \\
\hat{p} & -i \hbar \frac{\partial}{\partial x}
\end{aligned}
$$

Note: $\frac{\hbar}{i} \frac{\partial \psi}{\partial x}=E \psi$
Here $\psi$ is called an eigan function of the operator $-i \hbar \frac{\partial}{\partial x}$ and $E$ are called the corresponding energy eigan values.

## Schrodinger's wave equation:

Schrodinger's wave equations are the fundamental equations of quantum mechanics in the same sense as the Newton's second equation of motion of classical mechanics.

It is the differential form of de-Broglie wave associated with a particle and describes the motion of particle.


Figure(10): Schrodinger wave equations

## Schrodinger's time dependent wave equation in 1-dimentional form:

Let us assume that the $\psi$ for a particle moving freely in the positive x -direction is

$$
\begin{equation*}
\psi=A e^{-\frac{i}{\hbar}(E t-p x)} \tag{1}
\end{equation*}
$$

Now the total energy

$$
E=K E+(\text { Wotential energy })
$$

And we know that the kinetic energy is related with the momentum as $K E \quad \frac{p^{2}}{\overline{2} m}$
So the equation (2) in terms of wave function $\psi$, can be written as

$$
\begin{equation*}
E \psi=\left(\frac{p^{2}}{2 m}\right) \psi+1 \tag{3}
\end{equation*}
$$

As we know that the energy and momentum operators are given by $E \psi=\frac{d \psi}{d t}$ and $p \psi \quad \frac{\hbar}{\bar{i}} \frac{d \psi}{d x}$ Putting the values in equation (3) we get

$$
\begin{align*}
& i \hbar \frac{d \psi}{d t}=\frac{1}{2 m}\left(\frac{\hbar}{i} \frac{d}{d x}\right)^{2} \psi+V \psi \\
& i \hbar \frac{d \psi}{d t}=-\frac{\hbar^{2}}{2 m} \frac{d^{2} \psi}{d x^{2}}+V \psi \ldots \ldots \tag{4}
\end{align*}
$$

Schrodinger's time independent wave equation in 3-dimentional form:

$$
i \hbar \frac{d \psi}{d t}=\frac{\hbar^{2}}{2 m}\left(\frac{d^{2}}{d x^{2}}+\frac{d^{2}}{d y^{2}}+\frac{d^{2}}{d z^{2}}\right) \psi+V \psi
$$

But the Laplacian operator is given as $\left.\nabla^{2}=\frac{d^{2}}{d x^{2}}+\frac{d^{2}}{d y^{2}}+\frac{d^{2}}{d z^{2}}\right)$
So the above equation can be written as-

$$
\begin{equation*}
i \hbar \frac{d \psi}{d t}=\frac{\hbar^{2}}{2 m} \nabla^{2} \psi+V \tag{5}
\end{equation*}
$$

## Schrodinger's time independent wave equation in 1-dimention:

Again from wave function-

$$
\begin{aligned}
\psi & =A e^{-\frac{i}{\hbar}(E t-p x)} \\
\psi & =A e^{-\frac{i}{\hbar} E t} \cdot \frac{i}{\hbar} p x \\
\psi & =A e^{\frac{i}{\hbar} p x} \cdot \frac{-i}{e^{\hbar} E t} \\
\psi & =\psi_{0} e^{-\frac{i}{\hbar} E t}
\end{aligned}
$$

Where $\psi_{0}=\stackrel{i}{A}{ }_{A B} x$
Now differentiating partially with respect to $t$ we get

$$
\begin{equation*}
\frac{\partial \psi}{\partial t}=-\frac{i E}{\hbar} \psi_{0} e^{-\frac{i}{\hbar} E t} \tag{7}
\end{equation*}
$$

Now differentiating partially with respect to $x$ we get

Again

$$
\begin{align*}
\frac{\partial \psi}{\partial x} & =\frac{\partial \psi_{0}}{\partial x} e^{-\frac{i}{\hbar} E t} \\
\frac{\partial^{2} \psi}{\partial x^{2}} & =\frac{\partial^{2} \psi_{0}}{\partial x^{2}} e^{-\frac{i}{\hbar} E t} \tag{8}
\end{align*}
$$

Putting the value from (7) and (8) into equation (5) we get i.e. $i \hbar \frac{\partial \psi}{\partial x} \quad \stackrel{\hbar^{2}}{\overline{2} m} \nabla^{2} \psi+V \psi$

$$
\begin{align*}
& \left.i \hbar\left[-\frac{i E}{\hbar} \psi_{0} e^{-\frac{i}{\hbar} E t}\right]=-\frac{\hbar^{2}}{2 m}\left[\frac{\partial^{2} \psi_{0}}{\partial x^{2}} e^{-\frac{i}{\hbar} E t}\right]+V \not\right\} e^{-\frac{i}{\hbar} E t} \\
& E \psi_{0} e^{-\frac{i}{\hbar} E t}=e^{-\frac{i}{\hbar} E t}\left[\frac{\hbar^{2}}{2 m} \frac{\partial^{2} \psi_{0}}{\partial x^{2}}+V_{\nless \psi}\right] \\
& E \psi_{0}=\frac{\hbar^{2}}{2 m} \frac{\partial^{2} \psi_{0}}{\partial x^{2}}+V \psi \\
& E \psi_{0}-V \psi_{0}=\frac{\hbar^{2}}{2 m} \frac{\partial^{2} \psi_{0}}{\partial x^{2}} \\
& \left(\begin{array}{ll}
E & \forall
\end{array}\right) \psi_{0}=\frac{\hbar^{2}}{2 m} \frac{\partial^{2} \psi_{0}}{\partial x^{2}} \\
& \frac{2 m}{\hbar^{2}}\left(\begin{array}{ll}
E & \forall
\end{array}\right) \psi_{0}=\frac{\partial^{2} \psi_{0}}{\partial x^{2}} \\
& \frac{\partial^{2} \psi_{0}}{\partial x^{2}}-\frac{2 m}{\hbar^{2}}\left(\begin{array}{ll}
E & \forall
\end{array}\right) \psi_{0}=0 \tag{10}
\end{align*}
$$

Schrodinger's time independent wave equation in 3-dimention form:

$$
\begin{aligned}
\left(\frac{\partial^{2} \psi_{0}}{\partial x^{2}}+\frac{\partial^{2} \psi_{0}}{\partial y^{2}}+\frac{\partial^{2} \psi_{0}}{\partial z^{2}}\right) \frac{2 m}{\hbar^{2}}\left(E-V_{0}\right)_{3} & =0 \\
\nabla^{2} \psi_{0}-\frac{2 m}{\hbar^{2}}(E \quad \forall) \psi_{0} & =0
\end{aligned}
$$

Where $\nabla^{2}=\left(\frac{\partial^{2}}{\partial x^{2}}+\frac{\partial^{2}}{\partial y^{2}}+\frac{\partial^{2}}{\partial z^{2}}\right)$

## Application of Schrodinger's wave equation:

Energy level and wave function of a particle enclosed in one dimensional box of infinite height:
Let us consider the case of a particle of mass moving along x -axis between two rigid walls A and B at $x=$ 0 and $x=$. The potential energy $V$ of the particle is given as

$$
V=\left\{\begin{array}{l}
0 \text { for } 0<x<L \\
\infty \text { for } x \leq 0 \text { and } x \geq L
\end{array}\right.
$$

Within the box, the Schrodinger wave equation is given by

Let

$$
\begin{align*}
\frac{\partial^{2} \psi}{\partial x^{2}}-\frac{2 m}{\hbar^{2}} E \psi & =0 \quad[\because \quad V \quad=\ldots . .0 .  \tag{1}\\
k^{2} & =\frac{2 m E}{\hbar^{2}}  \tag{2}\\
\frac{\partial^{2} \psi}{\partial x^{2}}-4 \hbar \psi & =0 \quad \ldots \ldots \ldots \ldots \ldots . . . . . . . . . . . . \tag{3}
\end{align*}
$$

This is a second order differential equation and its solution is given by

$$
\begin{equation*}
\psi=A \sin k x+B \mathrm{c} \tag{4}
\end{equation*}
$$

Where $A$ and $B$ are constants, the value of these constants can be calculated by the boundary conditions. By first boundary condition if $x=0 \Rightarrow \psi=0$

Then by equation (4) we get

$$
\Rightarrow \begin{array}{rlllll}
0 & =A \sin k 0+B \cos k 0 \\
& B & =0
\end{array}
$$

For second boundary condition $\psi=$ at0x $=$ thlen by equation (4)

$$
\begin{aligned}
0 & =A \sin k L+0 . & \cos k L \\
\Rightarrow & A \sin & =0
\end{aligned}
$$

But $A \neq \operatorname{so} 0 \sin \quad k L=0 \Rightarrow k L=n \pi(n=0,1,2,3 \ldots \ldots \quad \ldots)$

As we know that

$$
\begin{equation*}
k^{2}=\frac{n^{2} \pi^{2}}{L^{2}} \tag{5}
\end{equation*}
$$

By equation (2) and (5)

$$
\frac{2 m E}{\hbar^{2}}=\frac{n^{2} \pi^{2}}{L^{2}}
$$

For representing the $n^{\text {th }}$ energy level replacing $E$ BY $E_{n}$ we have

$$
\begin{aligned}
\frac{2 m E_{n}}{\hbar^{2}} & =\frac{n^{2} \pi^{2}}{L^{2}} \\
E_{n} & =\frac{n^{2} \pi^{2}}{L^{2}} \stackrel{\hbar^{2}}{\stackrel{\hbar^{2}}{2}} \\
E_{n} & =\frac{n^{2} \pi^{2}}{L^{2}} \times \stackrel{h^{2}}{4 \pi^{2} 2 m}
\end{aligned}
$$

$$
\begin{equation*}
E_{n}=\frac{n^{2} h^{2}}{8 m L^{2}} \tag{6}
\end{equation*}
$$

It is clear from expression (6) that inside an infinitely deep potential well, the particle can have only discrete set of energy i.e. the energy of the particle is quantised. The discreet energies are given by

$$
\begin{aligned}
E_{1} & =\frac{h^{2}}{8 m L^{2}} \\
E_{2} & =\frac{4 h^{2}}{8 m L^{2}}=4_{1} E \\
E_{3} & =\frac{9 h^{2}}{8 m L^{2}}=9_{1} E \\
E_{4} & =\frac{16 h^{2}}{8 m L^{2}}=16_{1} E \\
. . & =. . . . \\
. . & =. . \quad . .
\end{aligned}
$$

The constant A of equation (4) can be obtained by applying the normalization condition i.e.

$$
\begin{align*}
& \int_{x=0}^{x=L}|\psi|^{2} d x=1 \\
& \int_{0}^{L}|A \quad \sin | k d x=1 \\
& A^{2} \int_{0}^{L} \sin ^{2} k x \quad d x=1 \\
& A^{2} \int_{0}^{L}\left(\frac{1-\cos 2 k)}{2}\right) d \lambda=1 \\
& \frac{A^{2}}{2} \int_{0}^{L}(1-\cos ) d d x x=1 \\
& \frac{A^{2}}{2}\left[\{x\}_{0}^{L}-\left\{\frac{\sin 2 k}{2 k}\right\}_{0}^{L}\right]=1 \\
& \frac{A^{2}}{2}\left[\left(L-0 \frac{1}{2 k}(\sin 2 k L-\operatorname{sir}=1\right.\right. \\
& \frac{A^{2}}{2}\left[L \frac{1}{2 k}\left(\sin \frac{2 n \pi}{L} L-\sin \right] \quad \because \quad k \frac{2 n \pi}{\overline{\bar{L}}}\right. \\
& \frac{A^{2}}{2}\left[L \frac{1}{2 k}(0-] C=1 \quad \because \sin 2 n \pi\right. \\
& \frac{A^{2}}{2} L=1 \\
& A=\sqrt{\frac{2}{L}} \tag{7}
\end{align*}
$$

Now the wave function will be given by

$$
\psi=\sqrt{\frac{2}{L}} \sin \frac{n \pi}{L} x
$$

$$
(n=1,2,3
$$

